

Improvement of auroral 3D reconstruction methods

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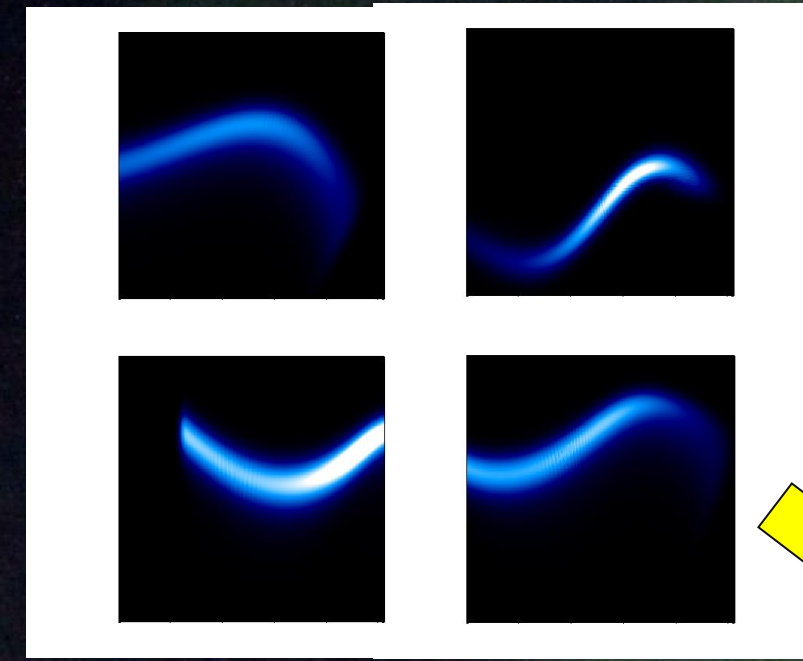
³ The Graduate University for Advanced Studies (SOKENDAI)

⁴ University of Electro-Communications

⁵ Kyushu Institute of Technology

Generalized – Aurora Computed Tomography

Multiple auroral images



Prior information

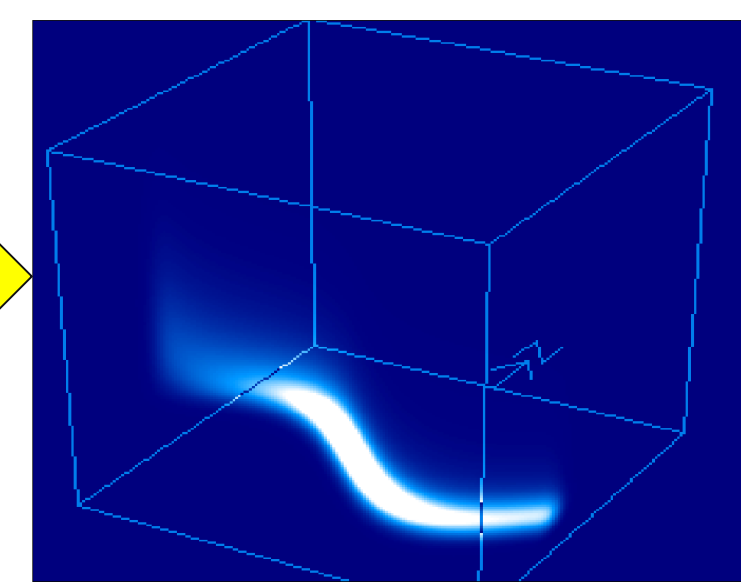
G-ACT is a method for reconstructing the 3D distribution of auroral emissions from **multi-instrument data**. [Tanaka et al., Ann. Geophys., 29, 551-562, 2011]

G - ACT

Bayesian inference

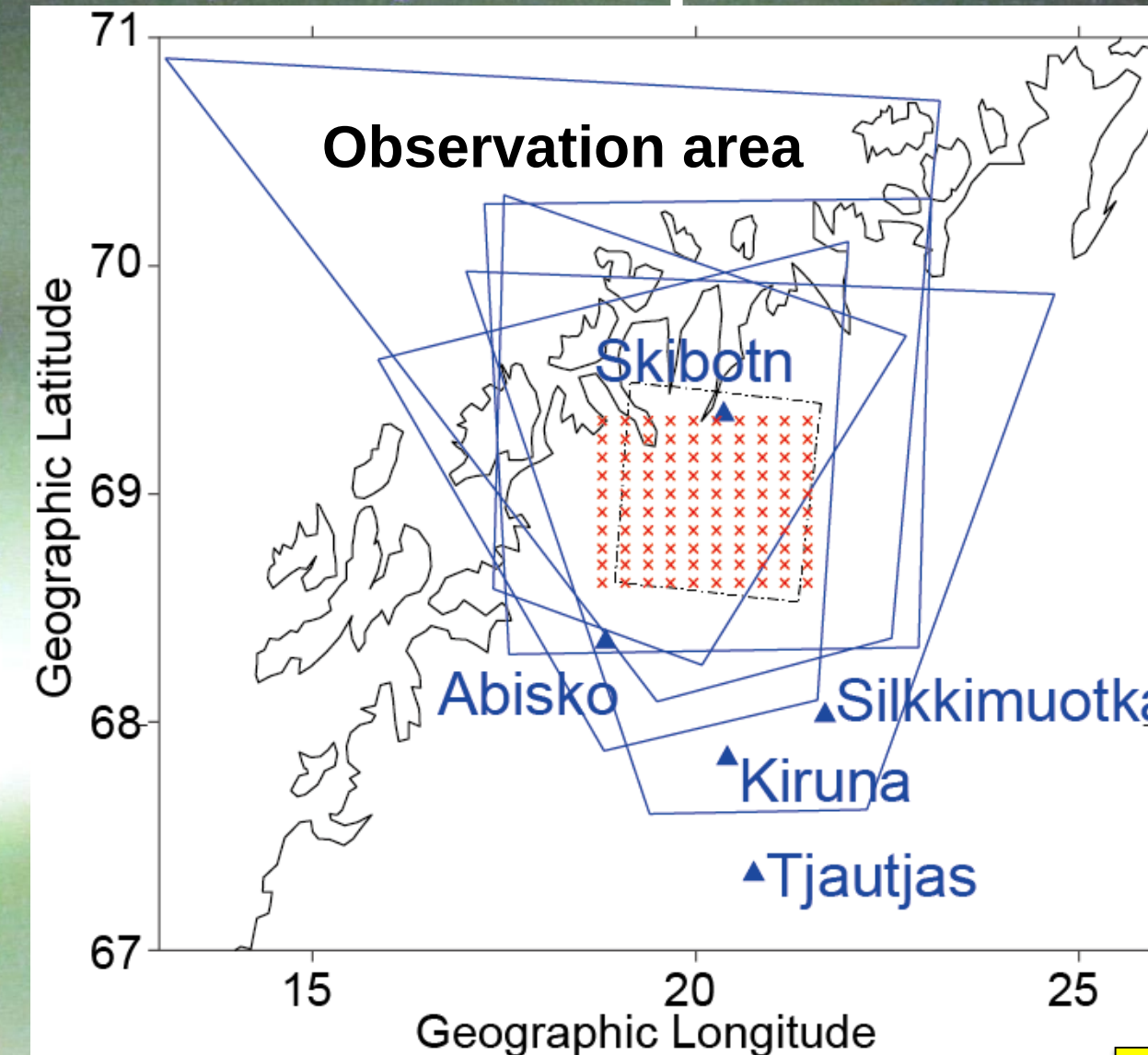
By using the 3D electron density data from **EISCAT_3D radar** for the G-ACT, the reconstruction results should be **more reliable** than that from the normal ACT.

3D distribution of electron density measured by EISCAT radar

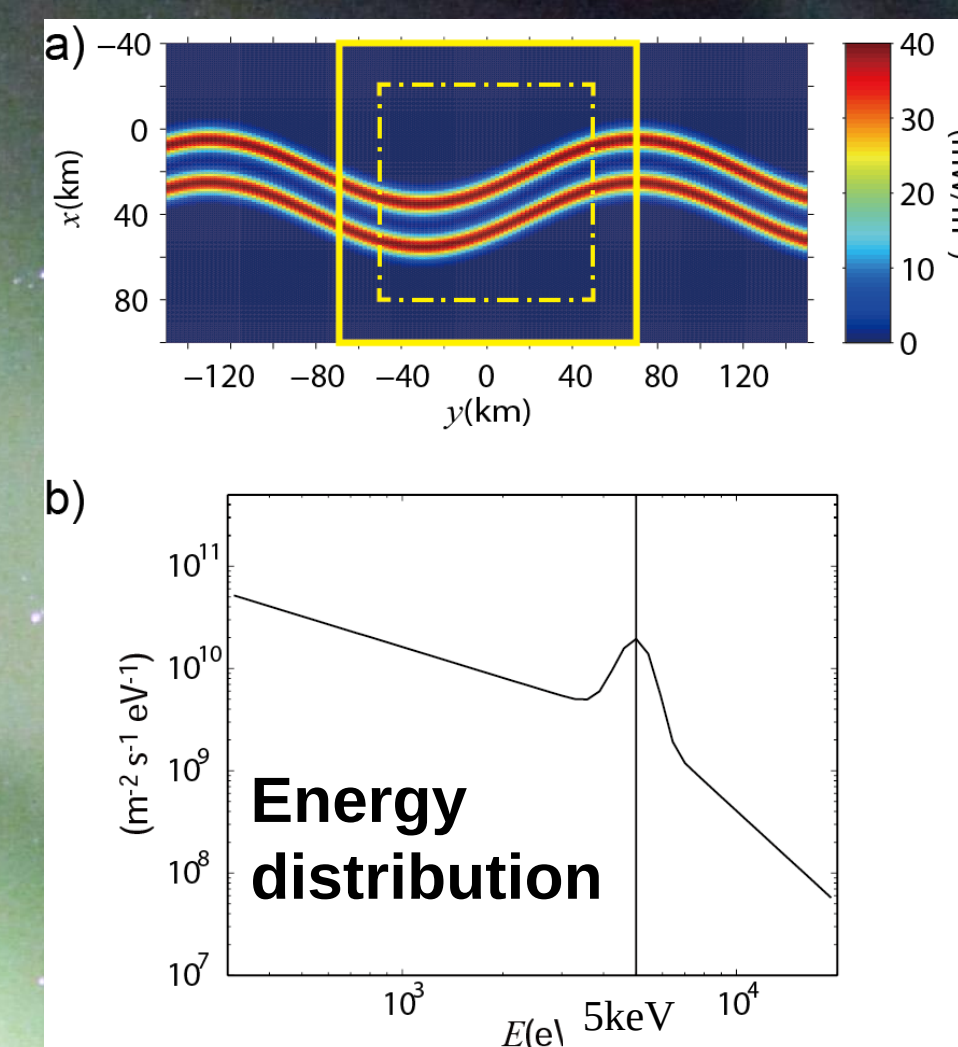


An example of G-ACT by simulation (Tanaka et al., 2024)

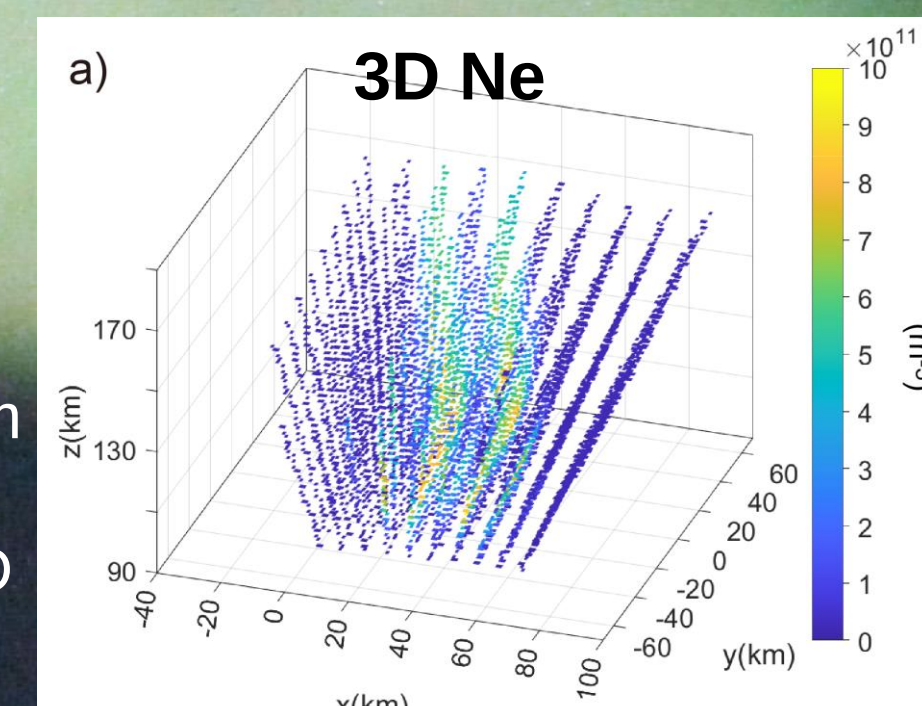
Assumptions



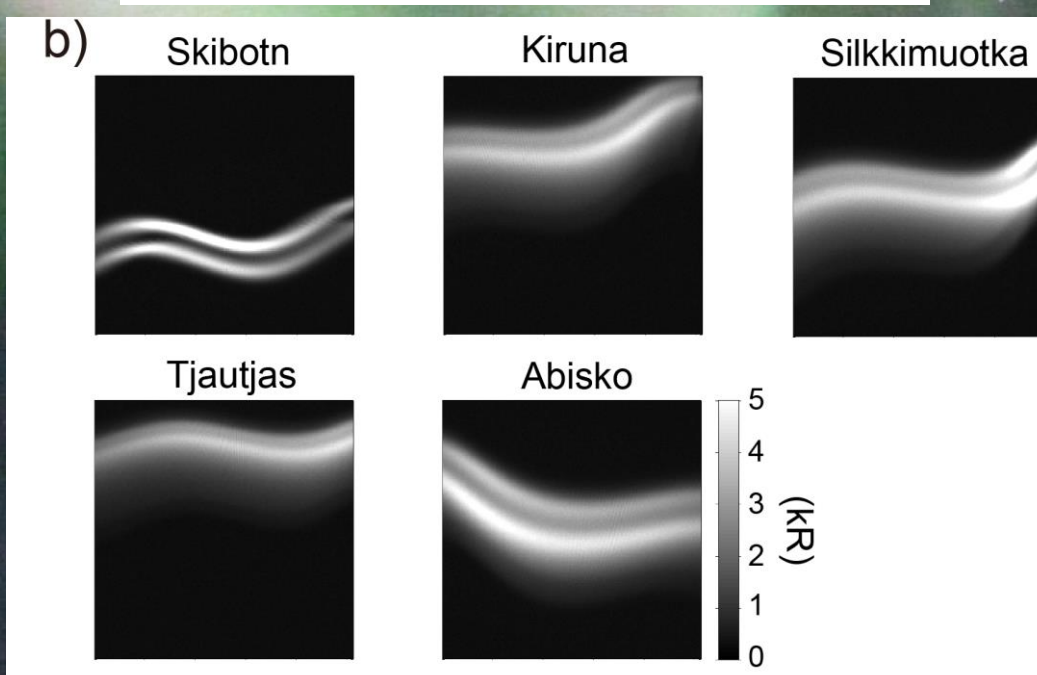
Auroral shape



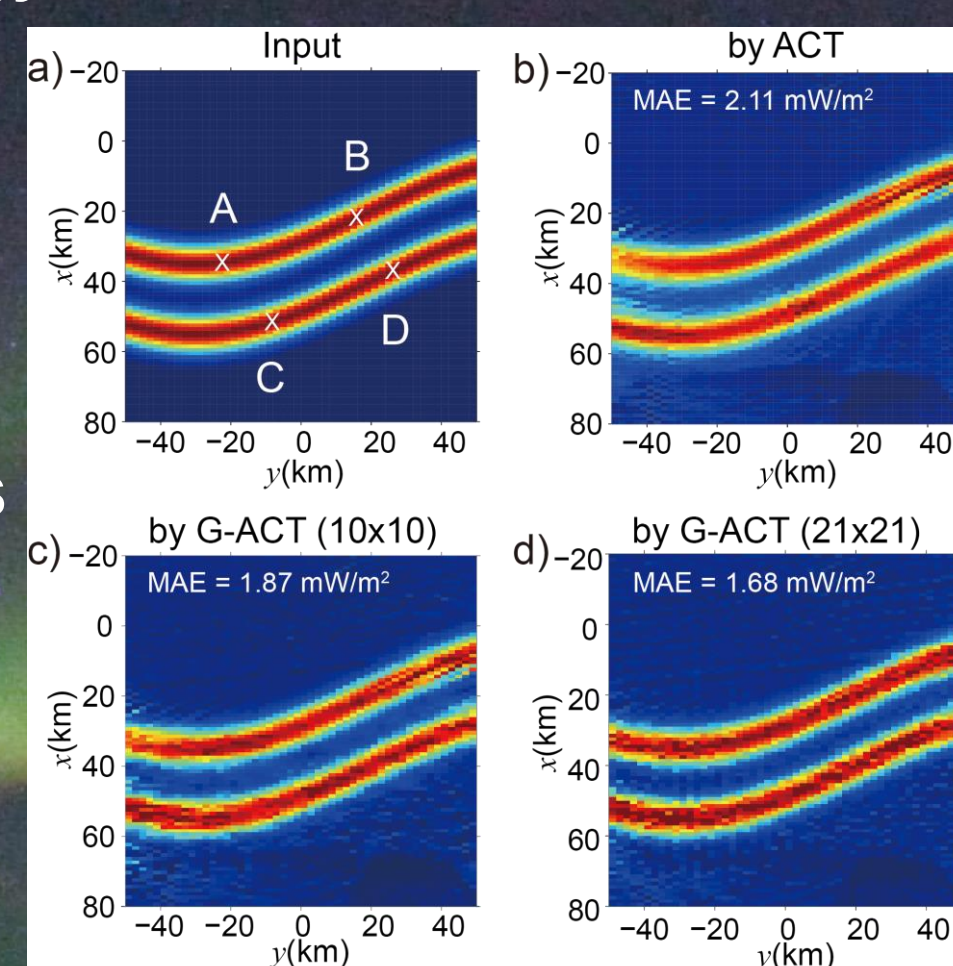
Simulated data



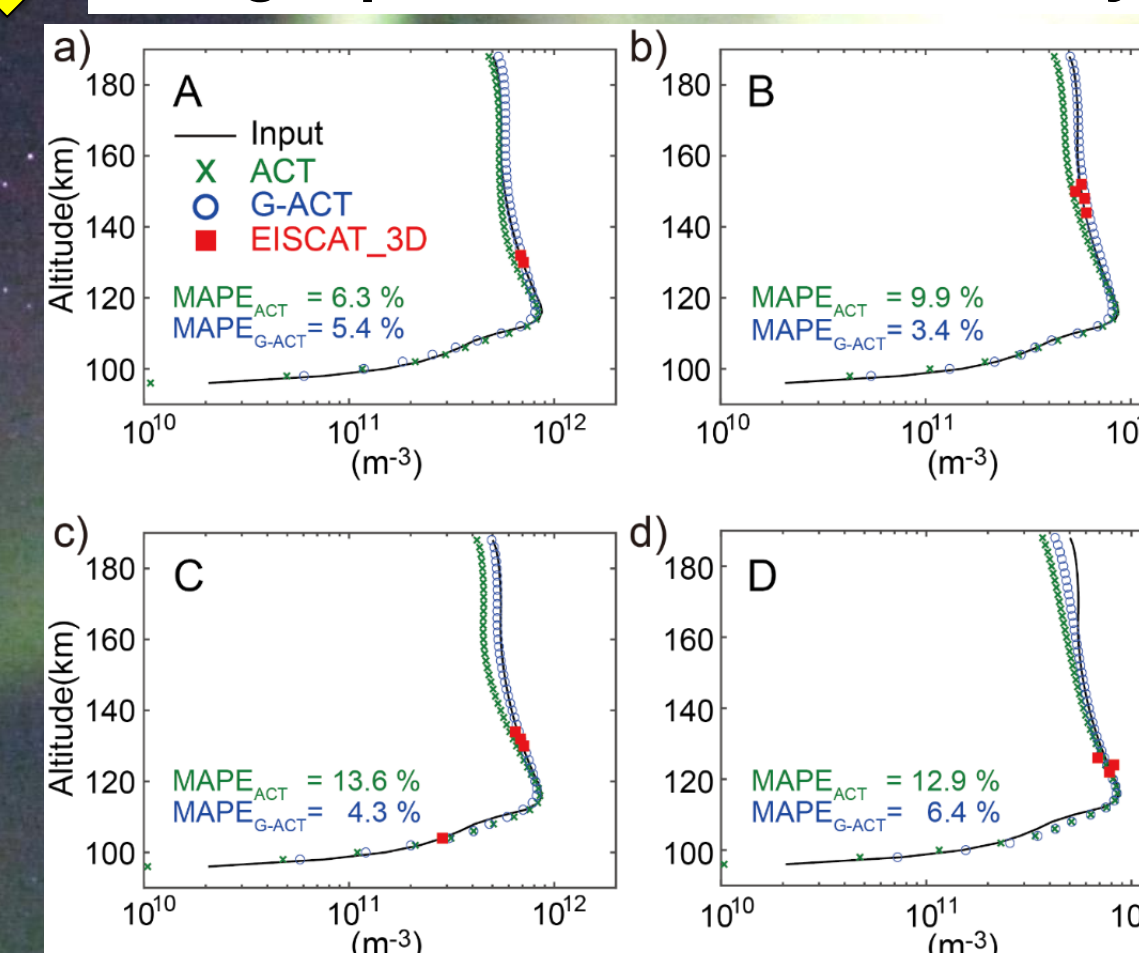
Images from 5 sites



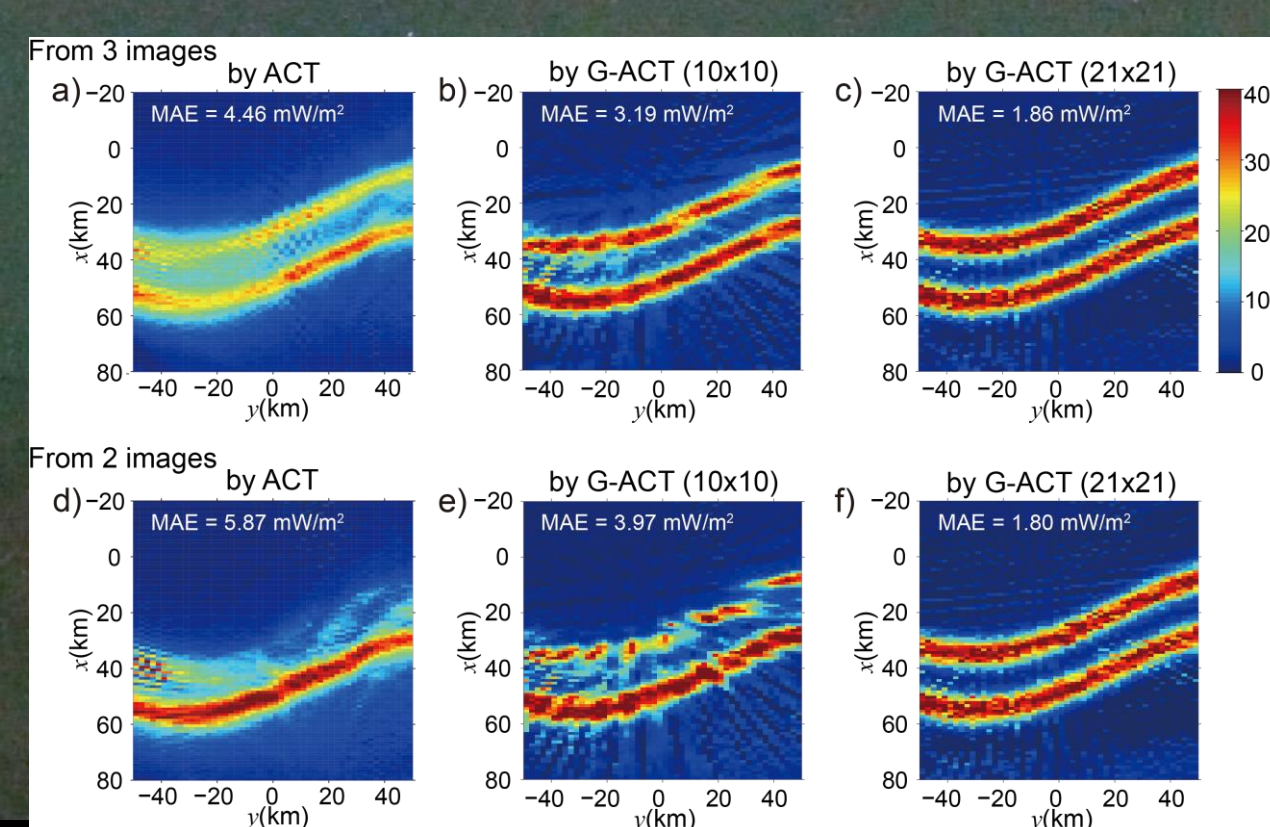
Reconstruction



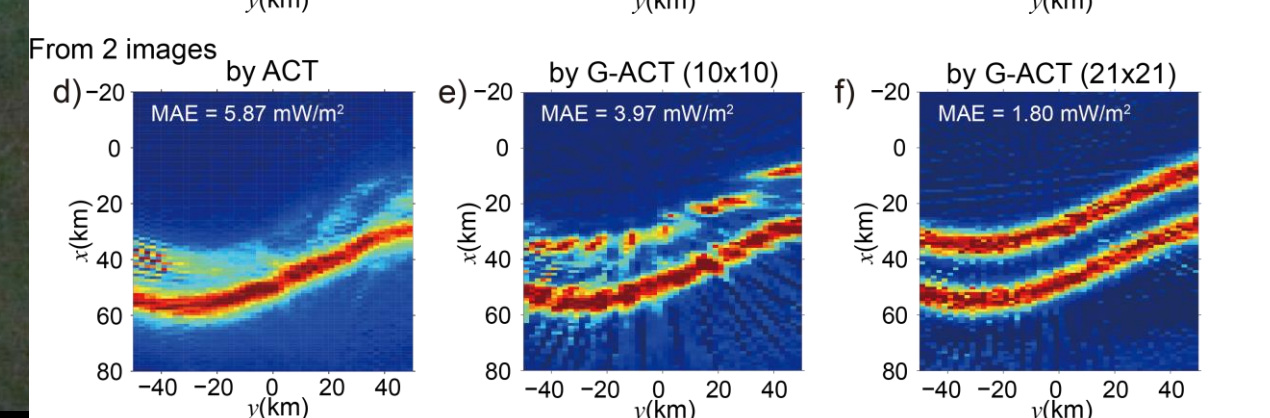
Height profile of electron density



From 3 images (S, K, T) + 3D Ne



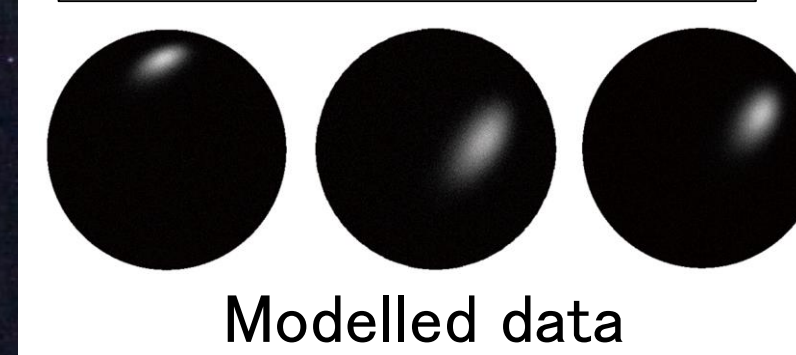
From 2 images (S, K) + 3D Ne



Improvement of analysis methods

Issues : It takes time to process the data and solve the inverse problem, in particular, for the **hyper-parameters**.

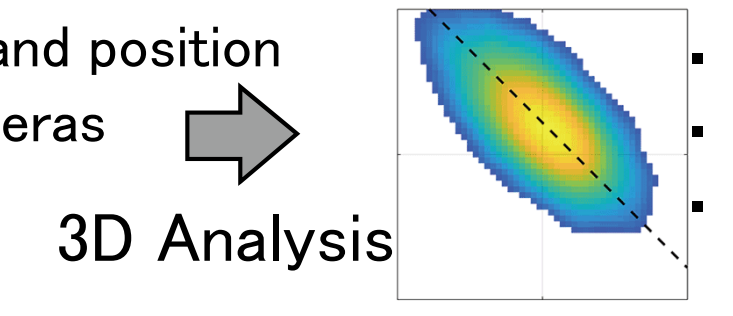
Simulation



Modelled data

Various observation conditions

- Auroral shape and position
- Number of cameras
- Wave length
- S/N ratio

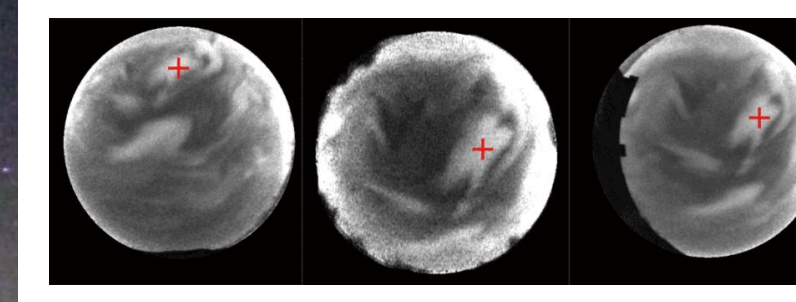


- Results
- Hyper-parameters
- Errors

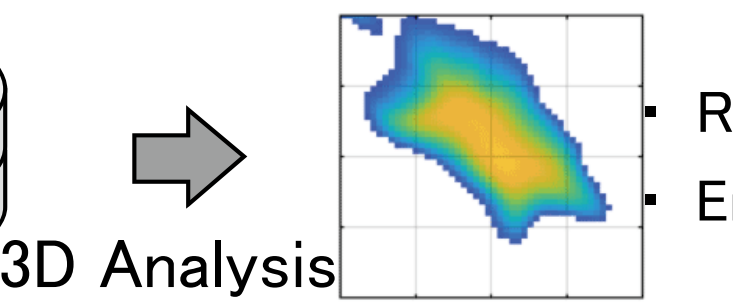
Database

Solution(1): The hyper-parameters, which are most time-consuming to determine, are **pre-calculated** under some typical conditions using model simulations and are compiled into a database to automatically determine the appropriate hyper-parameters for the observed data.

Observation



Observed data

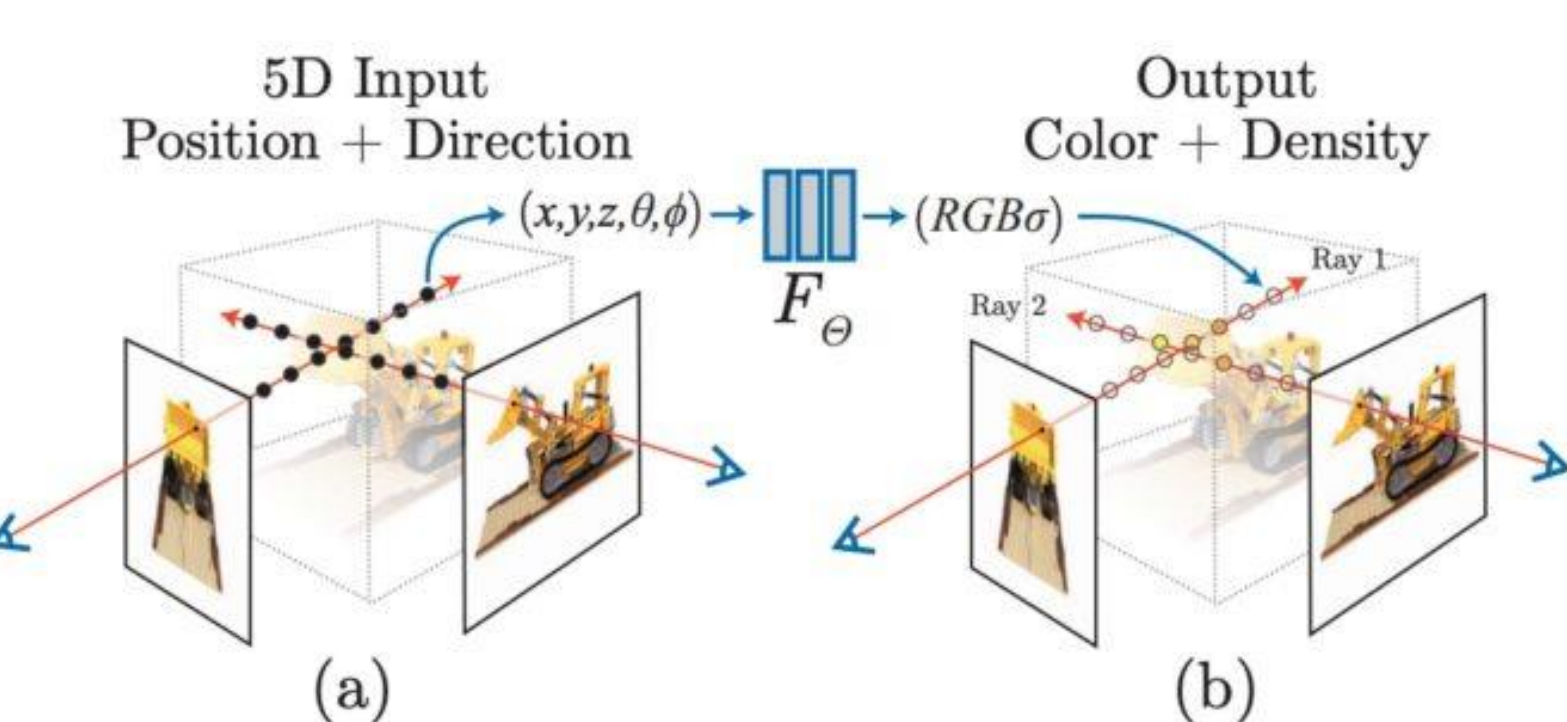


- Results
- Error evaluation

Solution(2): Use the state-of-the-art 3D reconstruction and visualization techniques based on AI, e.g., Neural Radiance Field (NeRF) or 3D Gaussian Splatting (3DGS).

- NeRF is a method based on deep learning for reconstructing a 3D representation of a scene from 2D images.

Neural Radiance Field (NeRF)



- We have started to develop the 3D reconstruction and visualization methods based on AI in collaboration with NII.

Summary

- To enhance the efficiency of 3D analysis, we are developing new procedures in which hyperparameters and errors are precomputed under typical conditions using model simulations and compiled into a database.
- We are also trying to develop AI-based 3D reconstruction and visualization methods in collaboration with NII. These methods can quickly and accurately reconstruct and visualize 3D scenes and have already been applied to various research fields.

Inverse analysis by the G-ACT method

Bayes' theorem: $P(\mathbf{f} | \tilde{\mathbf{b}}) = \frac{P(\tilde{\mathbf{b}} | \mathbf{f})P(\mathbf{f})}{P(\tilde{\mathbf{b}})} \propto P(\tilde{\mathbf{b}} | \mathbf{f})P(\mathbf{f}) \dots\dots(1)$

Likelihood: $P(\tilde{\mathbf{b}} | \mathbf{f}) \propto \exp\left\{-\sum_j \frac{1}{2}(\tilde{\mathbf{b}}_j - \mathbf{b}_j(\mathbf{f}))^T \Sigma_j^{-1}(\tilde{\mathbf{b}}_j - \mathbf{b}_j(\mathbf{f}))\right\}$ $P(\mathbf{f}) \propto \exp\left\{-\frac{\|\nabla^2 \mathbf{f}\|^2}{2\sigma^2}\right\}$

Posterior probability: $P(\mathbf{f} | \tilde{\mathbf{b}}) \propto \exp\left\{-\frac{1}{2\sigma^2}\left\{\sum_j w_j^2(\tilde{\mathbf{b}}_j - \mathbf{b}_j(\mathbf{f}))^T \Sigma_j^{-1}(\tilde{\mathbf{b}}_j - \mathbf{b}_j(\mathbf{f})) + \|\nabla^2 \mathbf{f}\|^2\right\}\right\} \dots\dots(2)$

$\mathbf{f}$: differential directional number flux of electrons
 j : type of data
 $\tilde{\mathbf{b}}$: observational data (auroral images, electron density)
 σ^2 : variance of $\nabla^2 \mathbf{f}$
 w_j : weighting factors for each data
 Σ_j^{-1} : inverse covariance matrix

$$\varphi(\mathbf{f}; w_j) = \sum_j w_j^2 (\tilde{\mathbf{b}}_j - \mathbf{b}_j(\mathbf{f}))^T \Sigma_j^{-1} (\tilde{\mathbf{b}}_j - \mathbf{b}_j(\mathbf{f})) + \|\nabla^2 \mathbf{f}\|^2$$

$$= w_1^2 \Sigma_{i,1}^{-1} \|\tilde{\mathbf{c}}_{i,1} - \mathbf{P}_{i,1} \mathbf{M}_1 \mathbf{f}\|^2 + w_2^2 \Sigma_2^{-1} \|\tilde{\mathbf{n}}_e - \mathbf{P}_2 \mathbf{M}_2 \mathbf{f}\|^2 + \|\nabla^2 \mathbf{f}\|^2 \dots\dots(3)$$

Hyper-parameters

Auroral images

Electron density